

## Abstract

*SuperNu*<sup>1</sup> is a Monte Carlo (MC) radiation transport code for simulating light curves of explosive outflows from supernovae. The MC transport step is domain replicated. To enable scaling on next-generation HPC systems, we are implementing the recursive coordinate bisection approach of domain decomposition for the opacity calculation. Then, we plan to propagate the decomposition to other steps in the simulation and construct a communication infrastructure to support the decomposition. In this poster we demonstrate the results of two communication schemes: the Improved KULL and Improved Milagro algorithms<sup>2</sup>.

## What is SuperNu?

Supernovae are stellar explosions, resulting in a cloud of gas called a nebula (Fig.1). The intent of SuperNu is to efficiently produce light curves and spectra for such nebulae. For example, Fig.2 present a light curve of a core-collapse supernova with a jet along one direction, modeled in 3D.

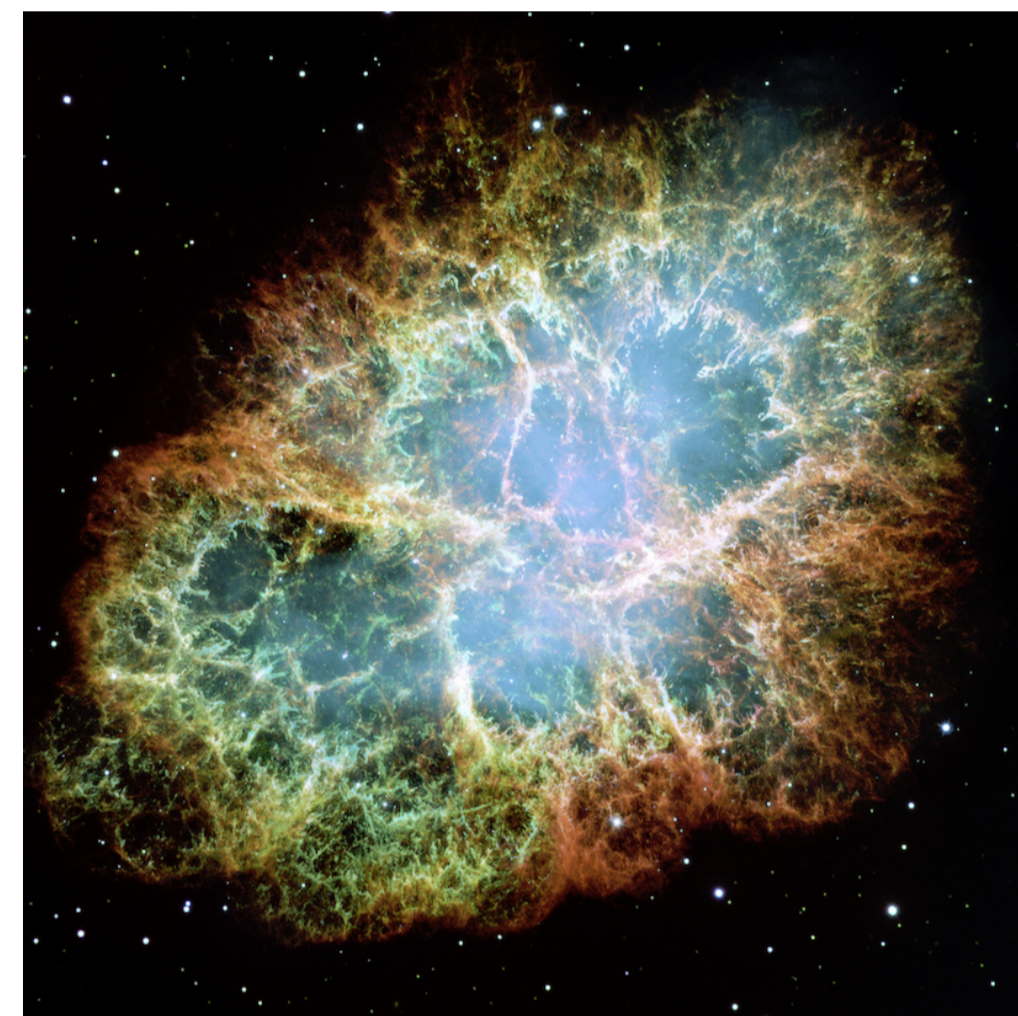


Figure 1: Crab Nebula, pictured above, was formed as result of a supernova<sup>4</sup>

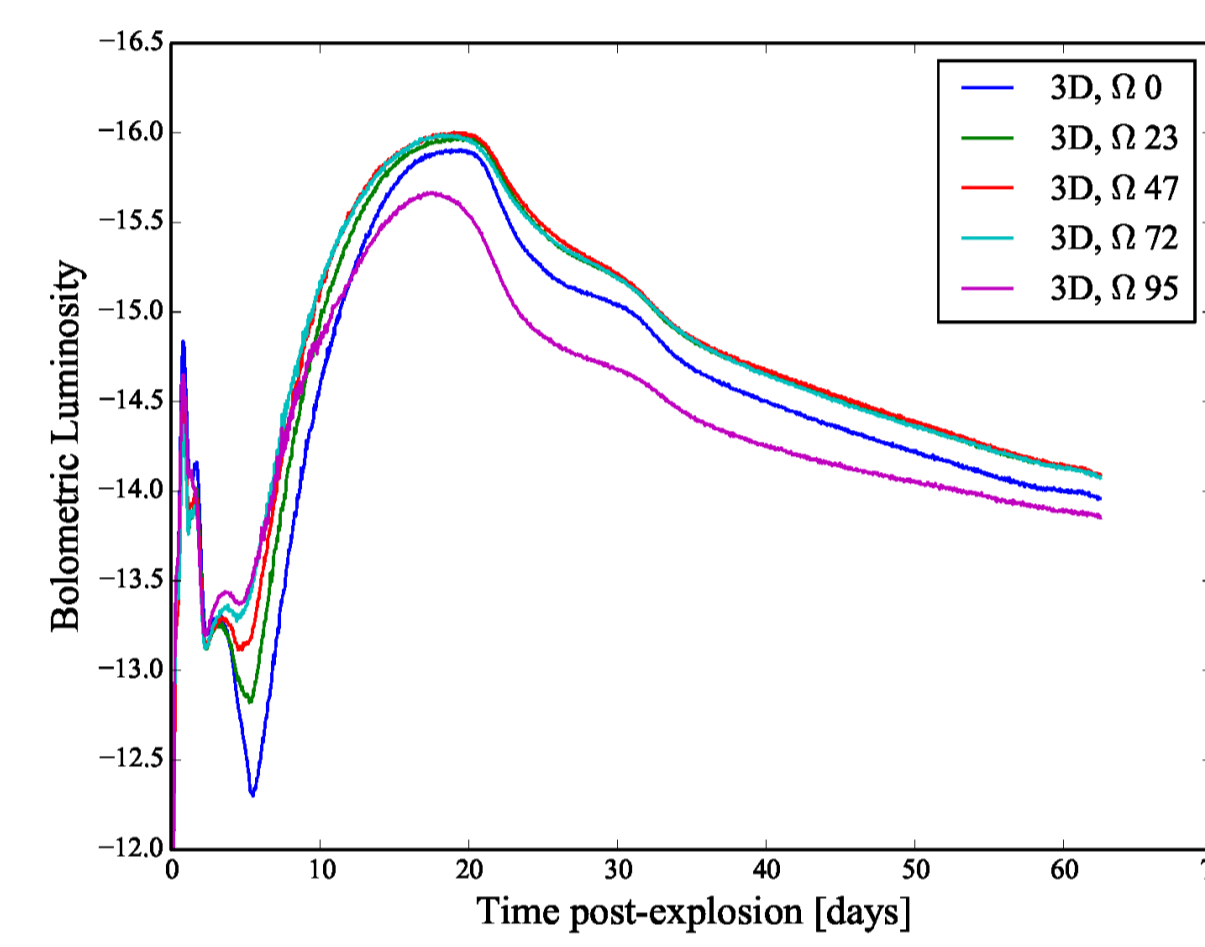


Figure 2: Light curve of a core-collapse supernova, where  $\Omega$  is the viewing angle

## Limitations

The previous particle transport approach involved replicating the whole domain on each rank, which limits problem size, memory wise.

The implementation also included an opacity calculation, where the domain was decomposed into strips (Fig.3). If this domain decomposition were applied to the transport step, there would likely be needlessly high communication between the ranks, as photons are less likely to stay in the part of the domain attributed to any one rank. This is due to the the low volume to surface area ratio of the strip decomposition.

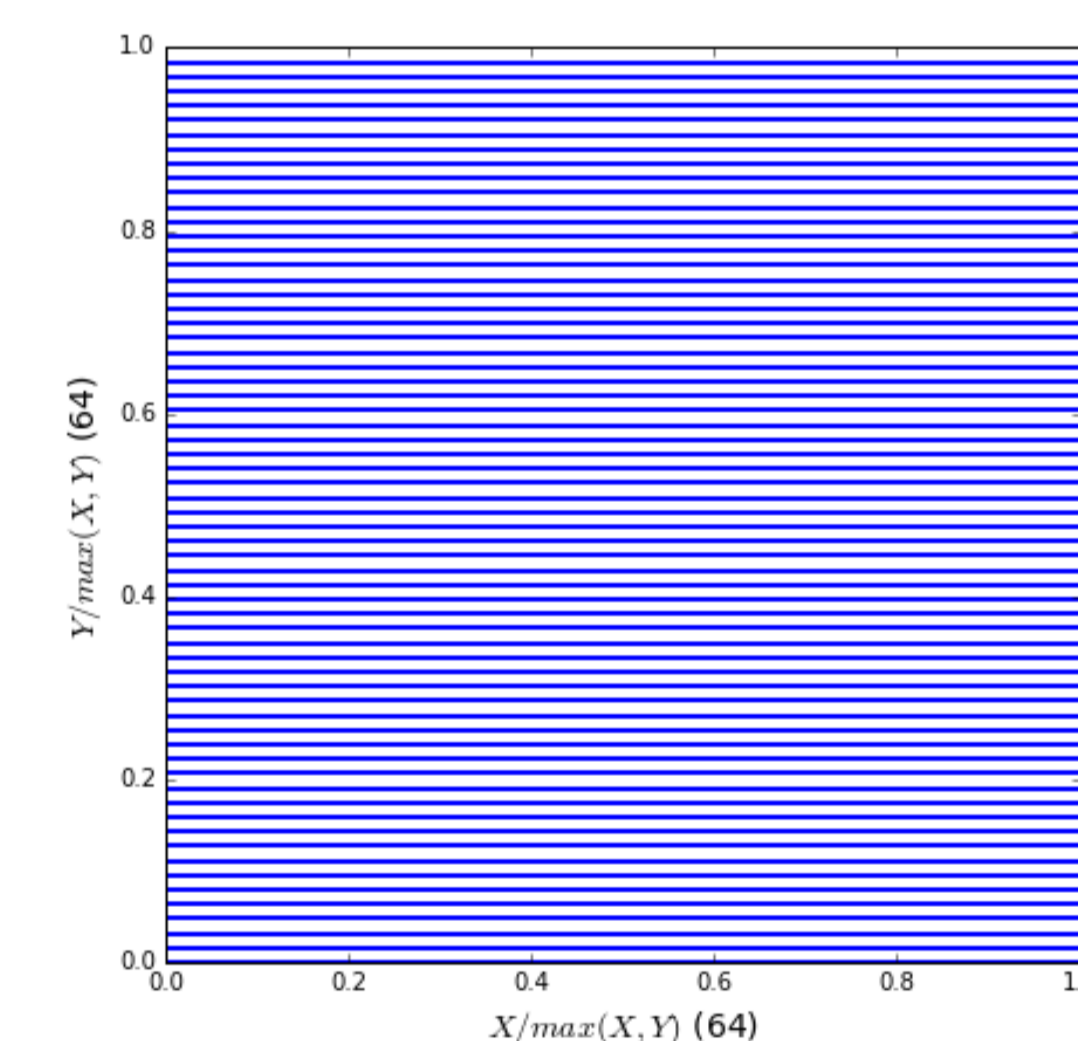


Figure 3: Particle sets in a strip configuration, with one strip of particles assigned to each of the 64 MPI rank.

## References

- 1) R.T. Wollaeger et al. 2013, ApJS 209, 37
- 2) T.A. Brunner et al, 2006, JCoPH .212..527B
- 3) R.T. Wollaeger, D.R. van Rossum. ApJ. 214 (2014) 28
- 4) J. Hester, A. Loll, NASA, ESA, (Arizona State University)

## Previous Performance

